





# Nuclear Structure in High and Low Energy Collisions

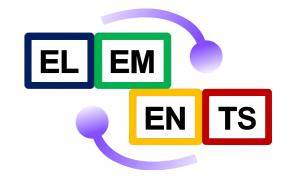
Hannah Elfner

January 26th, 2022, RBRC Workshop, BNL (virtual)









#### Outline

- High energy heavy-ion collisions
  - Deformation and neutron skin
  - Effects for isobar collisions
  - Nucleon-nucleon correlations
  - Color fluctuations
- Low beam energy collisions
  - SMASH with potentials
  - Collective anisotropic flow
  - Density changes and deformations
- Future plans
  - Short-range correlations in neutron-rich nuclei

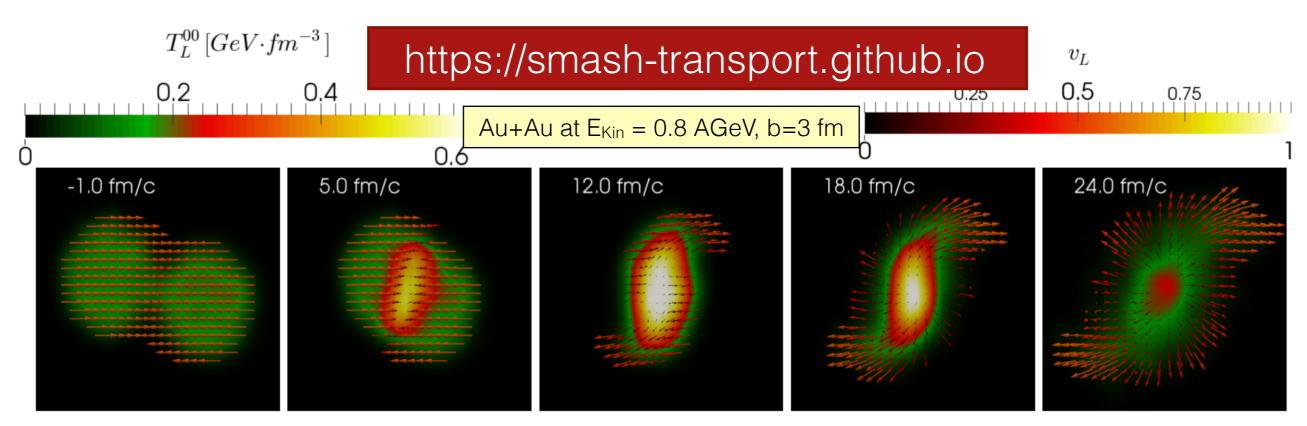


#### SMASH\*



Hadronic transport approach:

- J. Weil et al, PRC 94 (2016)
- Includes all mesons and baryons up to ~2 GeV
- Geometric collision criterion
- Binary interactions: Inelastic collisions through resonance/string excitation and decay
- Infrastructure: C++, Git, Doxygen, (ROOT)



\* Simulating Many Accelerated Strongly-Interacting Hadrons

#### The SMASH Team

- In Frankfurt:
  - Oscar Garcia-Montero
  - Gabriele Inghirami
  - Alessandro Sciarra
  - Jan Staudenmaier
  - Anna Schäfer
  - Justin Mohs
  - Jan Hammelmann
  - Niklas Götz
  - Renan Hirayama
  - Nils Saß
  - Antonio Bozic
  - Orhan Özel

- In US:
  - Dmytro Oliinychenko
  - Agnieszka Sorensen
- + Damjan Mitrovic



Group excursion in September 2020

### High Energy Collisions

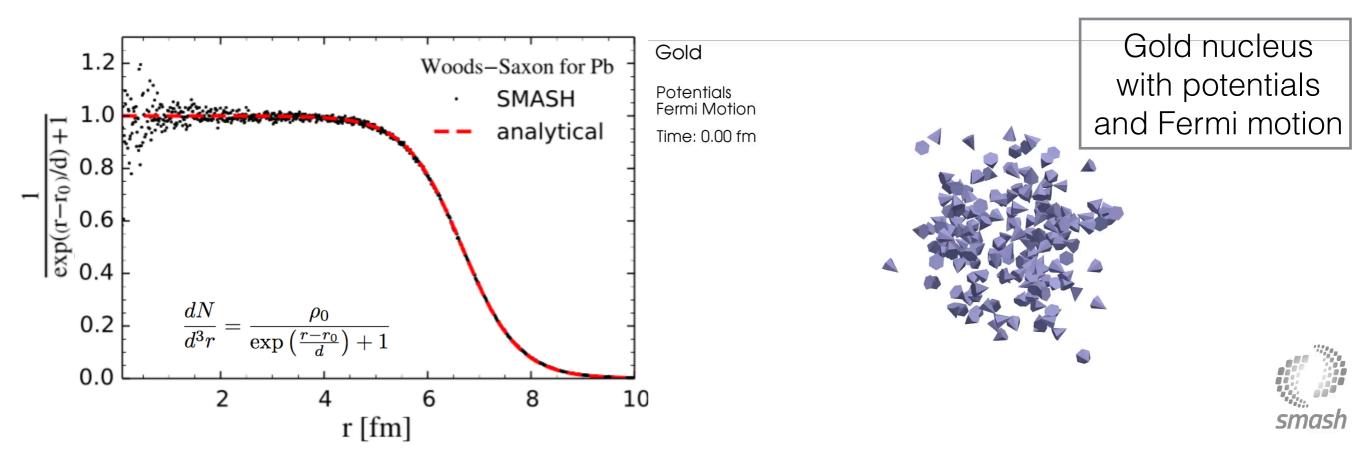
In collaboration with Alba Soto Ontoso, Massimiliano Alvioli, Mark Strikman

#### Initial Conditions

Nuclear Collisions

J. Weil et al, PRC 94 (2016)

Woods-Saxon distribution in coordinate space



- optional: deformed nuclei and (frozen) Fermi motion
- optional: read-in of more realistic initial states with correlations, neutron skin

#### Isobar Collisions

 Investigate potential maximal effect of deformation for Ru

$$\rho(r,\theta) = \frac{\rho_0}{e^{(r-R'(\theta,\phi))/d} + 1}$$

$$R'(\theta) = R_0(1 + \beta_2 Y_2^0(\theta)).$$

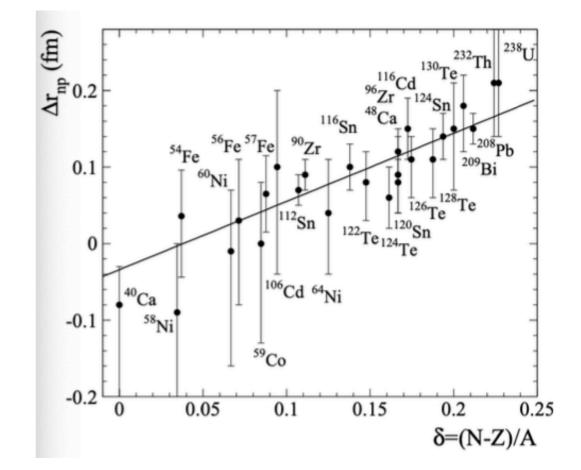
Nucleus	$R_0$ [fm]	d [fm]	$\beta_2$
$^{96}_{40}{ m Zr}$	5.02	0.46	0
96 44 Ru	5.085	0.46	0.158

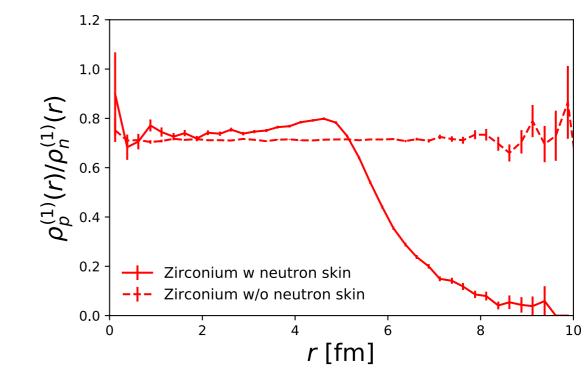
• And neutron skin for Zr, choice halo  $\Delta r_{np} = \langle r_n^2 \rangle^{1/2} - \langle r_p^2 \rangle^{1/2}$ 

$$\Delta r_{np}\Big|_{^{96}_{40}\mathrm{Zr}} = 0.12 \pm 0.03 \; \mathrm{fm}$$

Nucleon in $^{96}_{40}\mathrm{Zr}$	$R_0$ [fm]	d [fm]
p	5.08	0.34
n	5.08	0.46

J. Hammelmann et al, *Phys.Rev.C* 101 (2020)





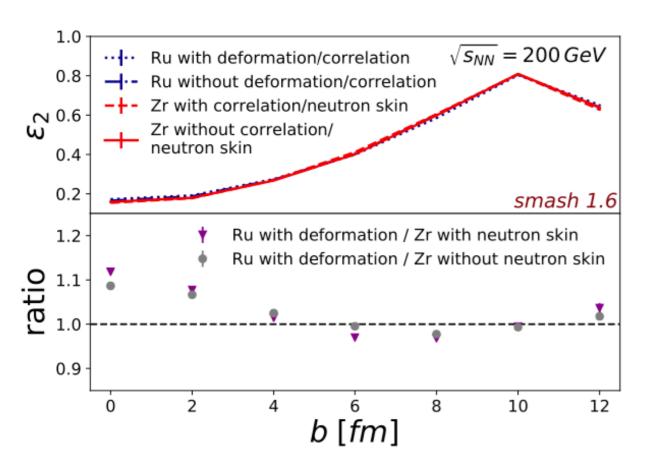
### Participant Eccentricity

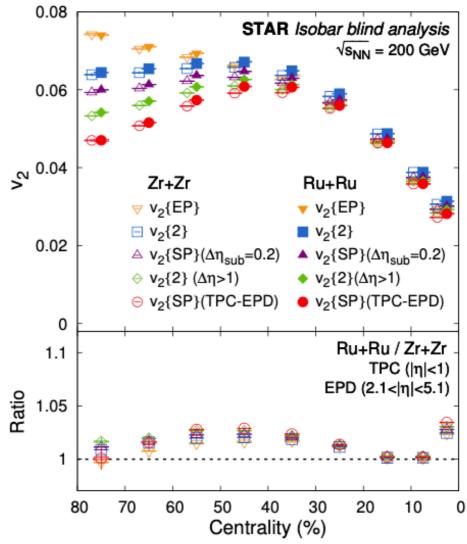
 Including nuclear structure effects and nucleon-nucleon correlations with initial state from full wave function

M. Alvioli, M. Strikman, PRC 100 (2019)

Hadronic transport approach SMASH is applied until full overlap

of nuclei



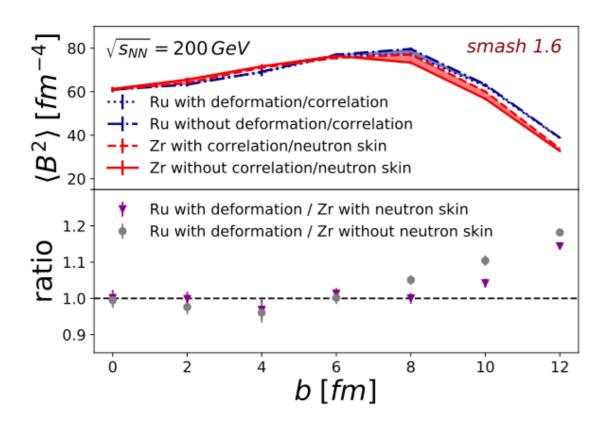


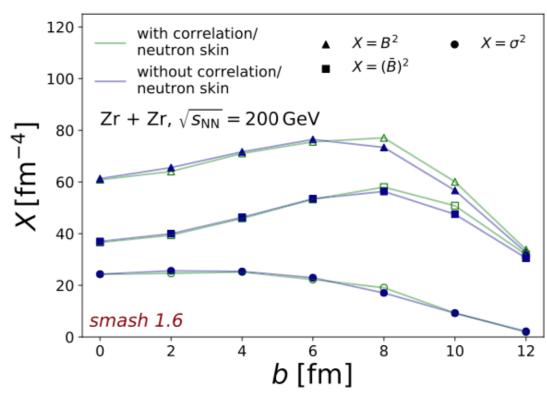
Participant eccentricity shows
 differences due to deformation at small impact parameters

J. Hammelmann et al, Phys.Rev.C 101 (2020) and STAR collaboration, Phys.Rev.C 105 (2022)

### Magnetic Field

 Due to the neutron skin, the charge is more concentrated in the middle -> differences in the magnetic field





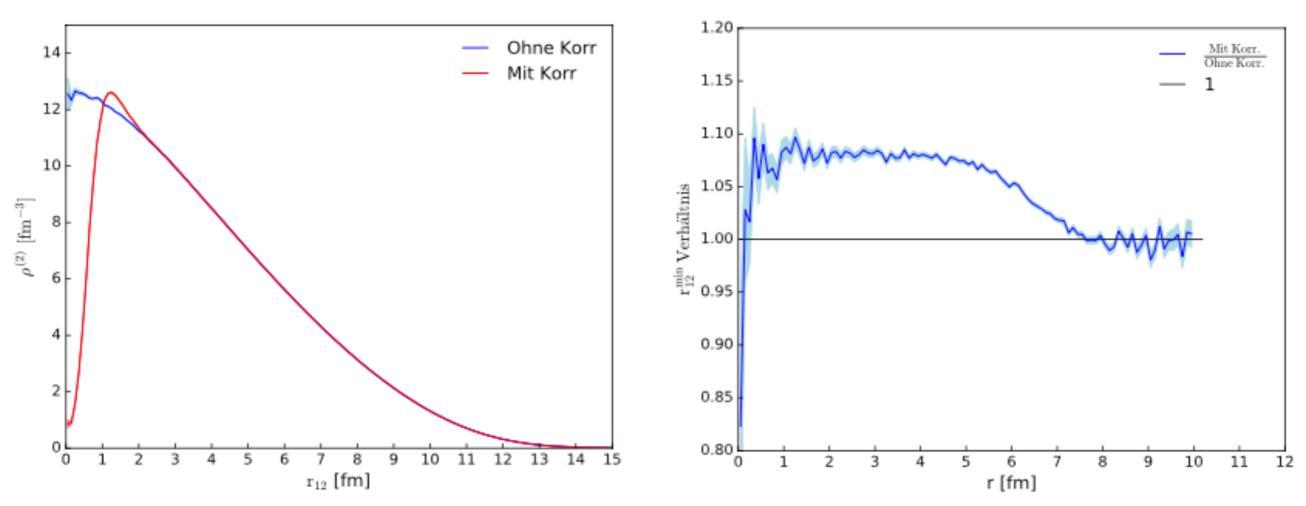
J. Hammelmann et al, Phys.Rev.C 101 (2020)

- The difference is really in the average field and not in the fluctuations
- One reason for missing difference between Ru/Zr results for CME correlators

  STAR collaboration, Phys. Rev. C 105 (2022)

#### NN correlations

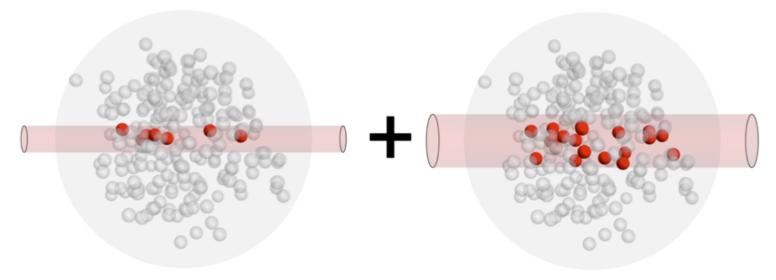
 Implementing nucleon nucleon correlations in the Au (and Cu) initial state in SMASH
 BSc thesis, Damjan Mitrovic, 2018



- The 2-particle distribution and the average distance shows the expected behaviour
- Other observables (eccentricities not sensitive)

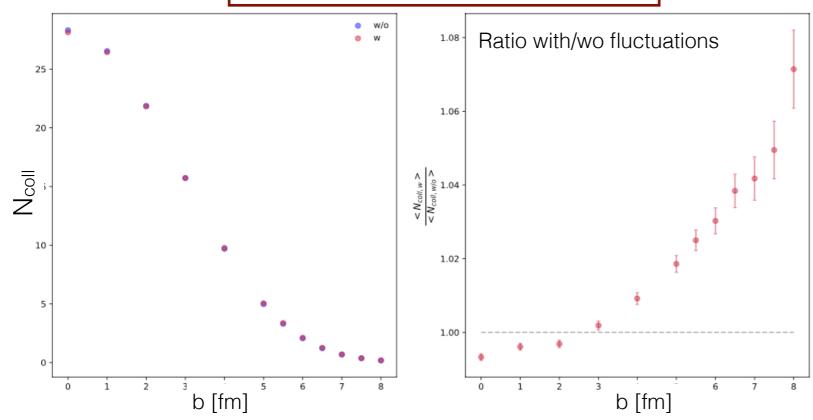
#### Color Fluctuations

#### M. Alvioli et al, PRD 98 (2018)



 Fluctuating first cross-section of NN interactions in SMASH





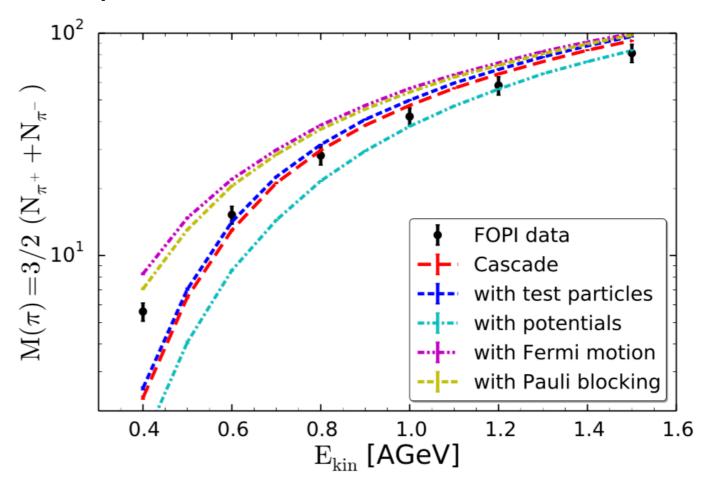
- Number of collisions in peripheral events is increased for small nuclei
- Other bulk observables are not affected

BSc thesis, Antonio Bozic, 2021

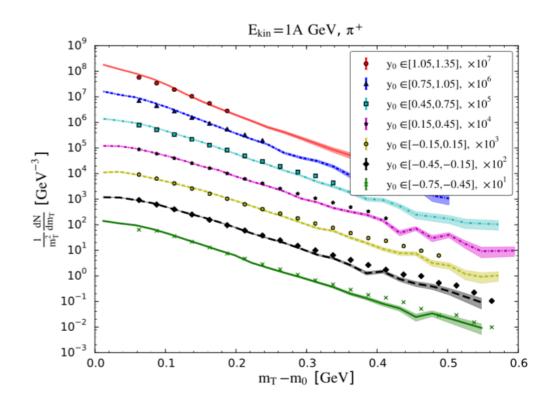
## Low Energy Collisions

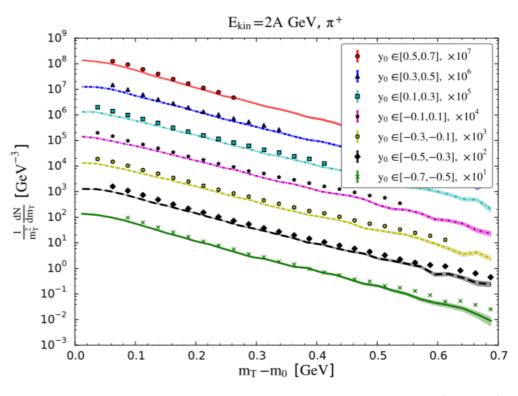
#### Pion Production in Au+Au

- Potentials decrease pion production, while Fermi motion increases yield
- Nice agreement with SIS experimental data



Note: consecutive addition of features

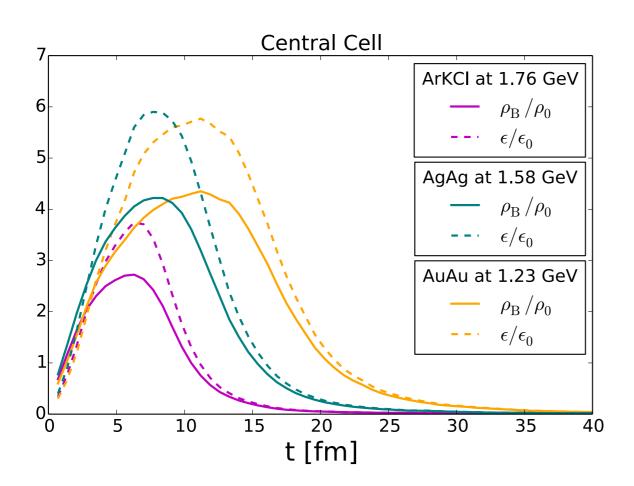


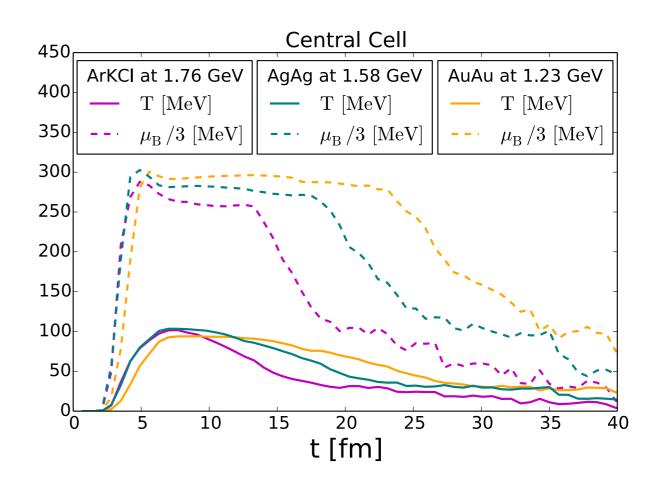


J. Weil et al, PRC 94 (2016)

#### Time Evolution

 Density and temperature in a central cell for heavy ion collisions at SIS-18 energies





J. Staudenmaier, N. Kübler and HE, arXiv:2008.05813

2-4 times nuclear ground state density reached

#### Collective Behaviour

- Potentials in SMASH
  - Basic Skyrme and symmetry potential

$$U_{\text{Skyrme}} = \alpha(\rho/\rho_0) + \beta(\rho/\rho_0)^{\tau}$$
  $U_{\text{Symmetry}} = \pm 2S_{\text{Pot}} \frac{\rho_{I_3}}{\rho_0}$ 

 Describes interactions between nucleons, repulsive at high densities

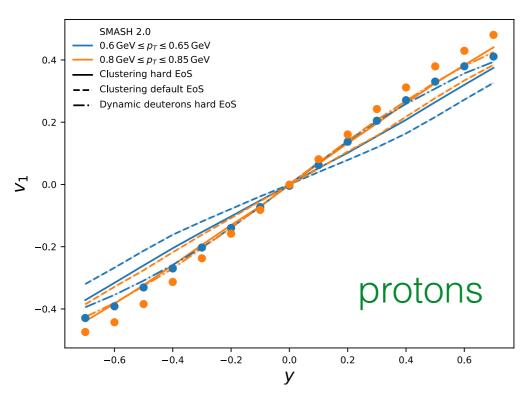
	soft EoS	default EoS	hard EoS
$\alpha$	$-356.0~\mathrm{MeV}$	$-209.2~\mathrm{MeV}$	$-124.0~{ m MeV}$
β	$303.0~{ m MeV}$	156.4 MeV	71.0 MeV
au	1.17	1.35	2.00
$\kappa$	200 MeV	240 MeV	380 MeV

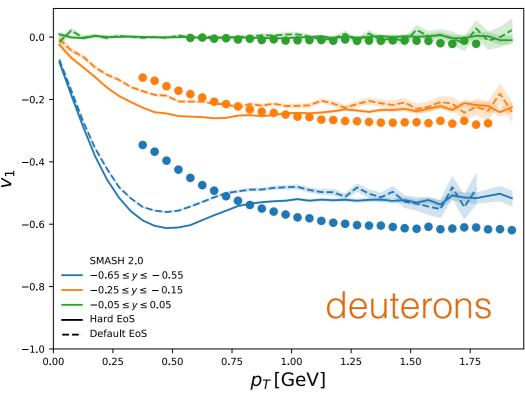
Default values according to transport code comparison

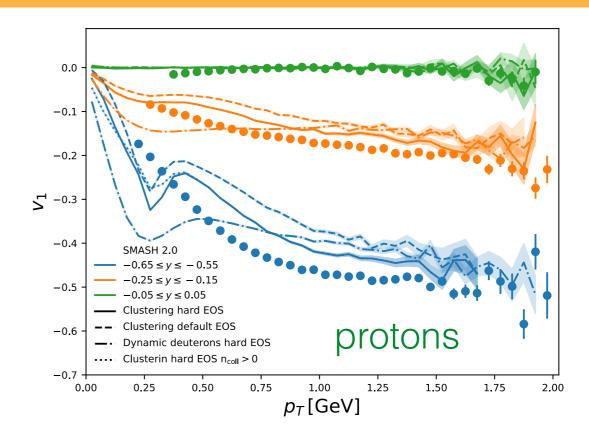
J. Xu et al., PRC 93 (2016)

#### Directed Flow in SMASH

#### J. Mohs, M. Ege, H. Elfner and M. Mayer, arXiv: 2012.11454

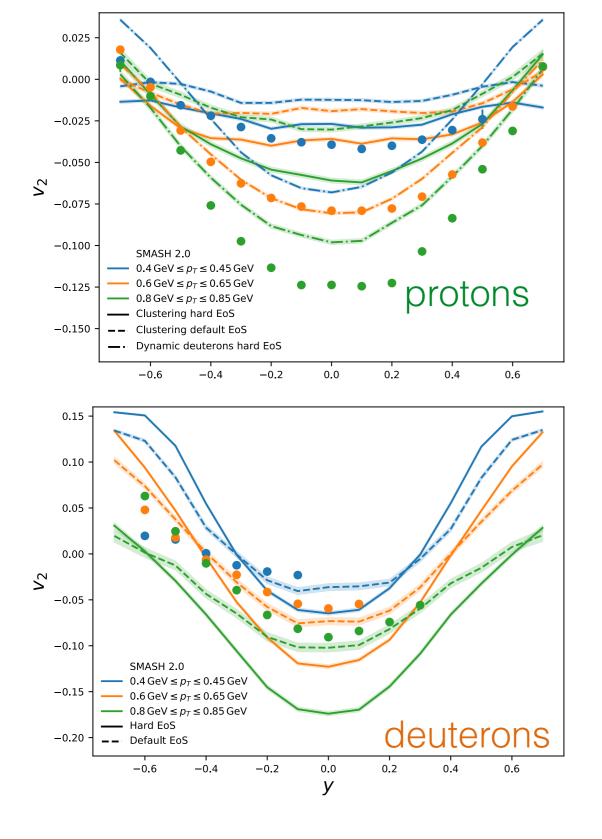


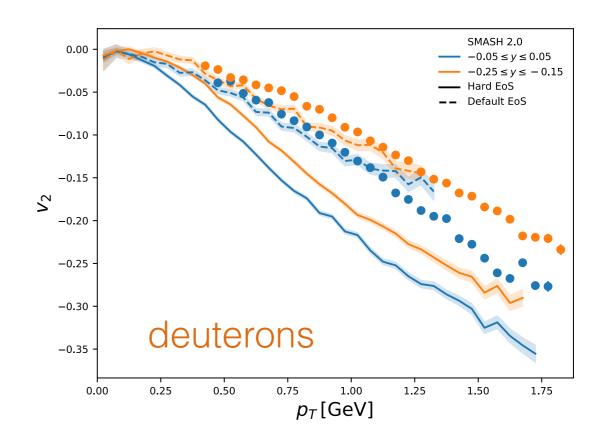




- Protons and deuterons fit better with hard EoS
- No momentum dependence of potential yet
- Clustering effect has similar magnitude as influence of potential

#### Elliptic Flow in SMASH

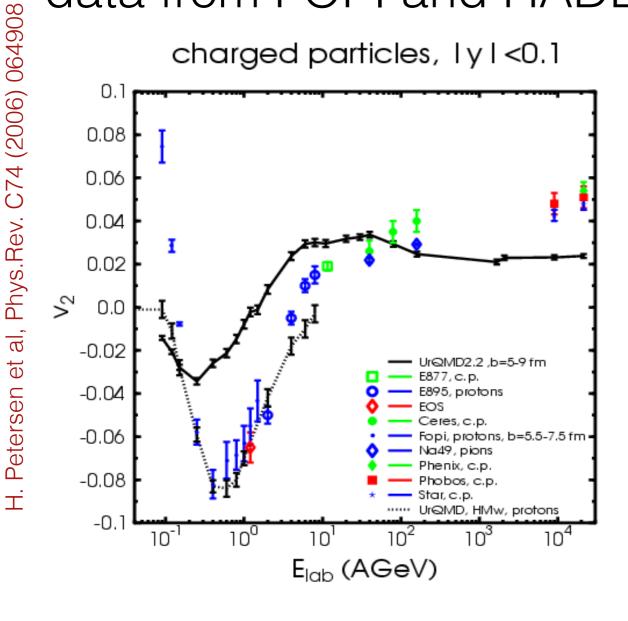


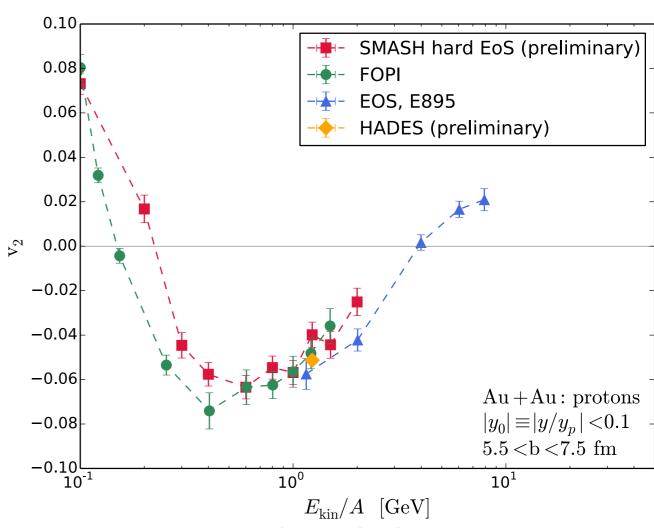


- Deuterons look better than protons (with default EoS)
- Again light clusters play a role and dependence on EoS is clearly visible

J. Mohs, M. Ege, H. Elfner and M. Mayer, arXiv: 2012.11454

 Directed and elliptic flow are compared to available data from FOPI and HADES





H.Petersen (now Elfner) et al, NPA 982, 2019

SMASH agrees well with previous UrQMD calculation

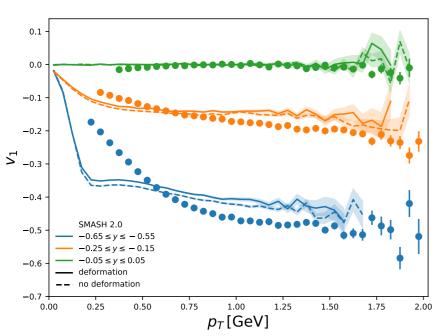
RBRC Workshop 01/26/2022

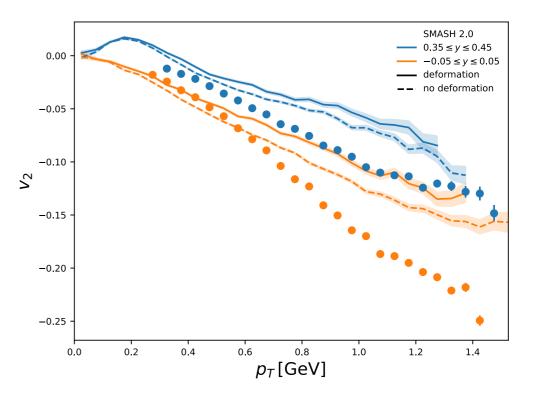
### Deformations and Density Effects

Deformations at low energy have some effect:

Artificially deformed Au nucleus to see qualitative

difference

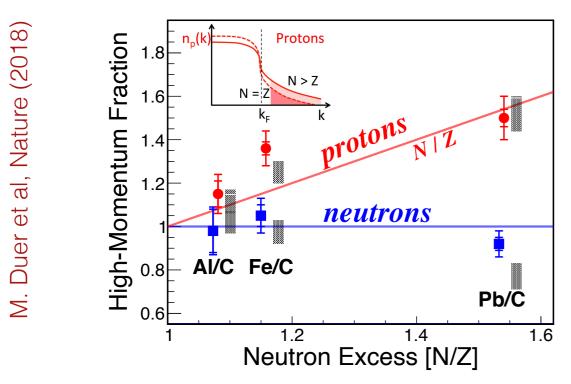


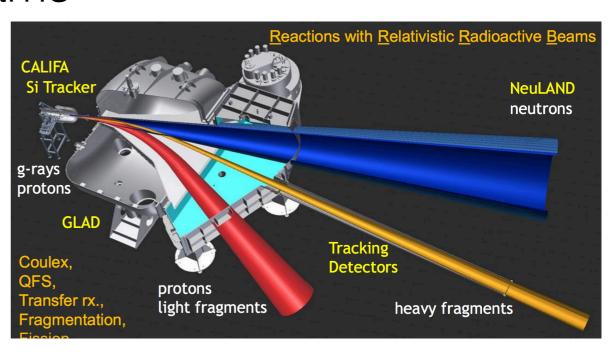


- Neutron skin has no effect on dilepton production in AuAu collisions at 1.23 AGeV
- Slight error in radius calculation within SMASH resulted in significant differences in the density calculation
  - Nuclear structure is important in low energy reactions

#### Future Plans

- Short-range correlations result in fluctuations to high momentum
- Develop a link between SRC observables and dense nuclear matter
- Identification of SRC in neutron-rich unstable nucleus at GSI with radioactive-ion beams





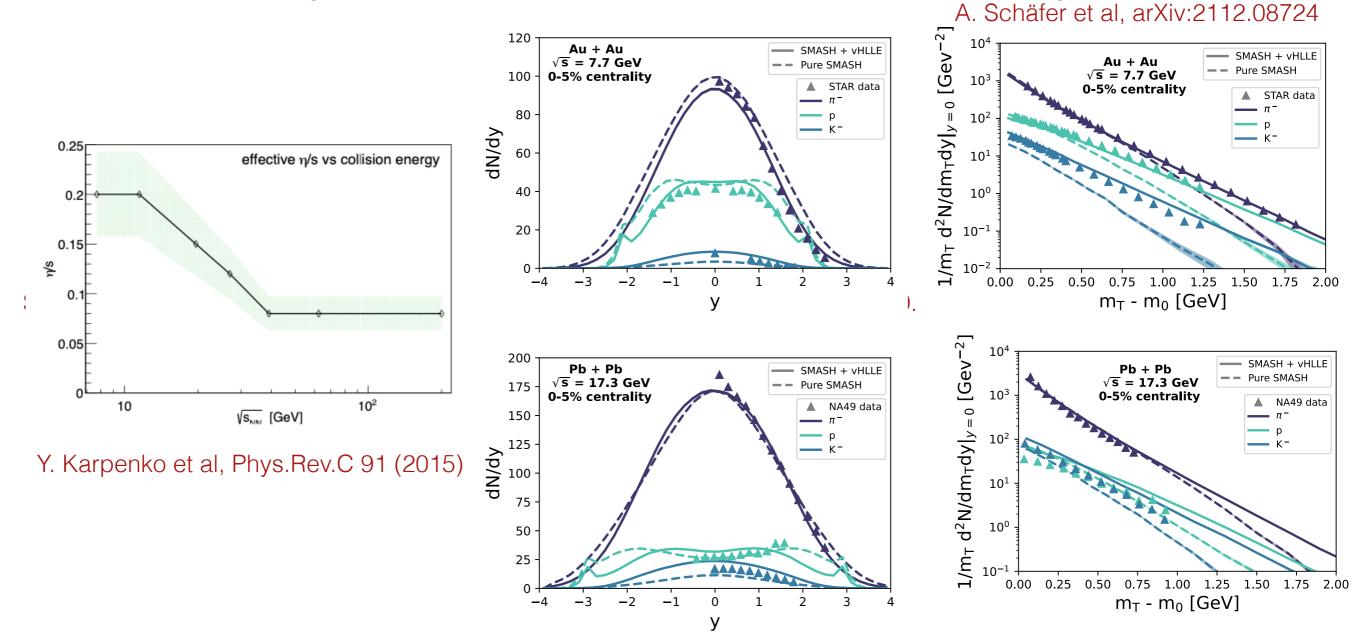
First experiments 16C+p vs 12C+p in 2022 (ELEMENTS)

### Summary

- High energy heavy ion collisions
  - SMASH employed for initial state evolution until full overlap
  - Deformation affects eccentricities in central collisions
  - Neutron skin affects magnetic field in peripheral collisions
  - NN correlations and color fluctuations play only minor role
- Low energy heavy-ion collisions
  - Collective flow is sensitive to nuclear mean field
  - Clustering in light nuclei is crucial
  - Deformation affects flow as well as changes in the density
- Source code is public and possibility to read in external initial state configurations is available
- Future: Short-range correlations with multi-GeV rare isotope collisions at FAIR

### SMASH Hybrid

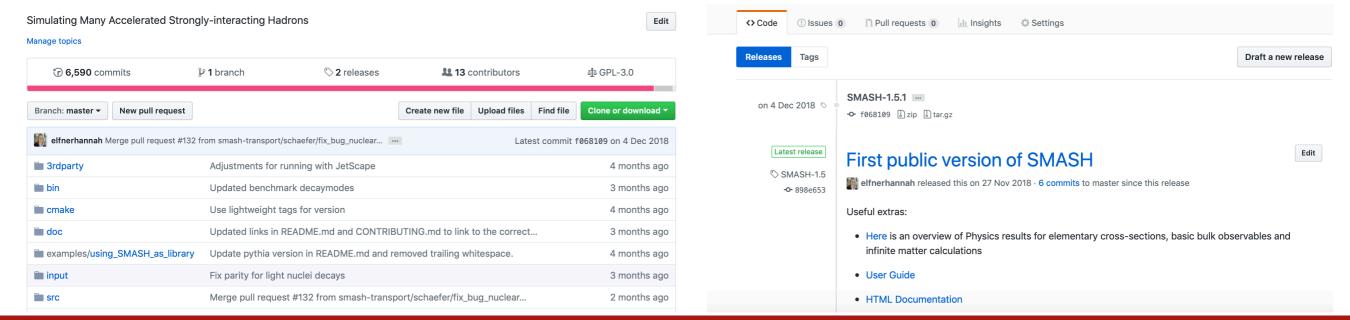
Full event-by-event SMASH+vHLLE+SMASH hybrid



See talk by A. Schäfer, Thursday, 9.30 AM (EST)

#### How to Use SMASH?

- Visit the webpage to find publications and link to SMASH-2.1 results <a href="https://smash-transport.github.io">https://smash-transport.github.io</a>
- Download the code at <a href="https://github.com/smash-transport/smash">https://github.com/smash-transport/smash</a>
- SMASH-2.1 has HepMC and RIVET
- Checkout the Analysis Suite at <a href="https://github.com/smash-transport/smash-analysis">https://github.com/smash-transport/smash-analysis</a>
- Find user guide and documentation at <a href="https://github.com/smash-transport/smash/releases">https://github.com/smash-transport/smash/releases</a>
- Animations and Visualization Tutorial under <a href="https://smash-transport.github.io/movies.html">https://smash-transport.github.io/movies.html</a>



# Backup

### General Setup

 Transport models provide an effective solution of the relativistic Boltzmann equation

$$p^{\mu}\partial_{\mu}f_{i}(x,p) + m_{i}F^{\alpha}\partial_{\alpha}^{p}f_{i}(x,p) = C_{\text{coll}}^{i}$$

- Particles represented by Gaussian wave packets for density calculations
- Geometric collision criterion

$$d_{\rm trans} < d_{\rm int} = \sqrt{\frac{\sigma_{\rm tot}}{\pi}}$$

Test particle method

$$d_{\text{trans}}^2 = (\vec{r_a} - \vec{r_b})^2 - \frac{((\vec{r_a} - \vec{r_b}) \cdot (\vec{p_a} - \vec{p_b}))^2}{(\vec{p_a} - \vec{p_b})^2}$$

As in UrQMD

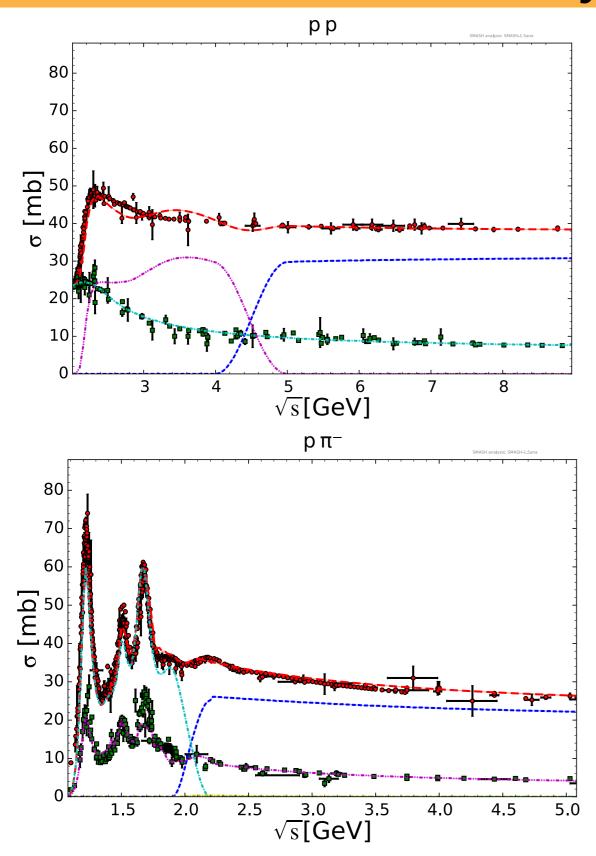
$$\sigma \mapsto \sigma \cdot N_{\text{test}}^{-1}$$
 $N \mapsto N \cdot N_{\text{test}}$ 

### Degrees of Freedom

N	Δ	٨	Σ	Ξ	Ω		Un	flavored		Strange
N <sub>938</sub>	Δ <sub>1232</sub>	۸1116	Σ <sub>1189</sub>	Ξ <sub>1321</sub>	Ω- <sub>1672</sub>	π <sub>138</sub>	f <sub>0 980</sub>	f <sub>2 1275</sub>	π <sub>2 1670</sub>	K <sub>494</sub>
$N_{1440}$	$\Delta_{1620}$	$\Lambda_{1405}$	$\Sigma_{1385}$	= <sub>1530</sub>	$\Omega$ -2250	$\pi_{1300}$	f <sub>0 1370</sub>	$f_{2'1525}$		K* <sub>892</sub>
$N_{1520}$	$\Delta_{1700}$	$\Lambda_{1520}$	$\Sigma_{1660}$	=1690		$\pi_{1800}$	f <sub>0 1500</sub>	<b>f</b> <sub>2 1950</sub>	<b>ρ</b> <sub>3 1690</sub>	K <sub>1 1270</sub>
$N_{1535}$	$\Delta_{1900}$	$\Lambda_{1600}$	Σ <sub>1670</sub>	Ξ <sub>1820</sub>			f <sub>0 1710</sub>	$f_{22010}$		K <sub>1 1400</sub>
$N_{1650}$	$\Delta_{1905}$	$\Lambda_{1670}$	Σ <sub>1750</sub>	= <sub>1950</sub>		η <sub>548</sub>		f <sub>2 2300</sub>	Фз 1850	K* <sub>1410</sub>
N <sub>1675</sub>	$\Delta_{1910}$	$\Lambda_{1690}$	Σ <sub>1775</sub>	Ξ <sub>2030</sub>		η΄ <sub>958</sub>	<b>a</b> <sub>0 980</sub>	f <sub>2 2340</sub>		$K_0^*_{1430}$
N <sub>1680</sub>	$\Delta_{1920}$	$\Lambda_{1800}$	Σ <sub>1915</sub>			η <sub>1295</sub>	<b>a</b> <sub>0 1450</sub>		<b>a</b> <sub>4 2040</sub>	$K_{2}^{*}_{1430}$
N <sub>1700</sub>	$\Delta_{1930}$	$\Lambda_{1810}$	Σ <sub>1940</sub>			η <sub>1405</sub>		f <sub>1 1285</sub>		K* <sub>1680</sub>
N <sub>1710</sub>	$\Delta_{1950}$	$\Lambda_{1820}$	$\Sigma_{2030}$			η <sub>1475</sub>	Ф1019	f <sub>1 1420</sub>	f <sub>4 2050</sub>	K <sub>2 1770</sub>
N <sub>1720</sub>		$\Lambda_{1830}$	$\Sigma_{2250}$				Ф1680			$K_{3}^{*}_{1780}$
N <sub>1875</sub>		$\Lambda_{1890}$				σ <sub>800</sub>		a <sub>2 1320</sub>		K <sub>2 1820</sub>
N <sub>1900</sub>		$\Lambda_{2100}$					h <sub>1 1170</sub>			$K_4^*_{2045}$
N <sub>1990</sub> N <sub>2060</sub>		$\Lambda_{2110}$				ρ <sub>776</sub>		$\pi_{11400}$		
N <sub>2080</sub>		$\Lambda_{2350}$				ρ <sub>1450</sub>	<b>b</b> <sub>1 1235</sub>	$\pi_{11600}$		<b>)</b>
N <sub>2100</sub>						ρ <sub>1700</sub>				
N <sub>2120</sub>							<b>a</b> <sub>1 1260</sub>	η <sub>2 1645</sub>		
N <sub>2190</sub>						ω <sub>783</sub>				
N <sub>2220</sub>						ω <sub>1420</sub>		ω <sub>3 1670</sub>		
N <sub>2250</sub>				Α	s of SMASH-1.7	ω <sub>1650</sub>				

- Mesons and baryons according to particle data group
- Isospin multiplets and anti-particles are included

### Elementary Cross Sections



- Total cross section for pp/pπ collisions
- Parameterized elastic cross section
- Many resonance contributions to inelastic cross section
- Reasonable description of experimental data
- Soft strings a la UrQMD and hard strings via Pythia 8

J. Weil et al, PRC 94 (2016), updated SMASH-1.5